

# Winds and Outflows in WiFeS and SALT + BIRD Kinematics

Moses Mogotsi

South African Astronomical Observator

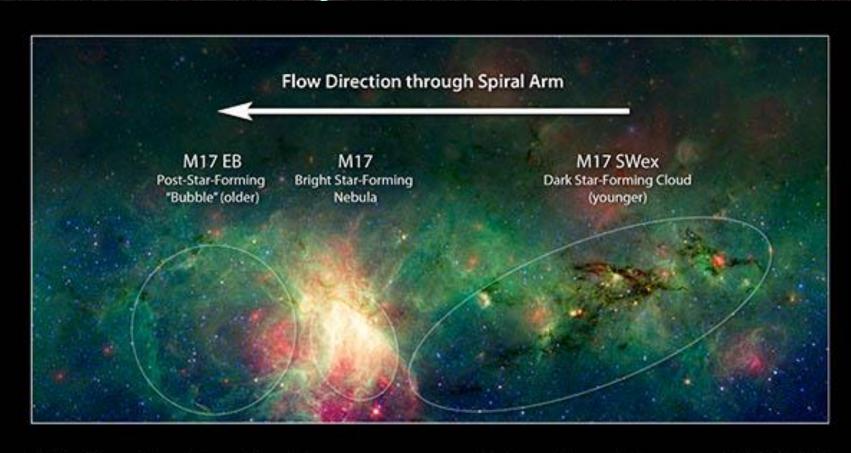


Moses Mogotsi, SUNBIRD, Cape Town, 20/11/2017

#### Star Formation & Baryon Cycle: Multi-wavelength

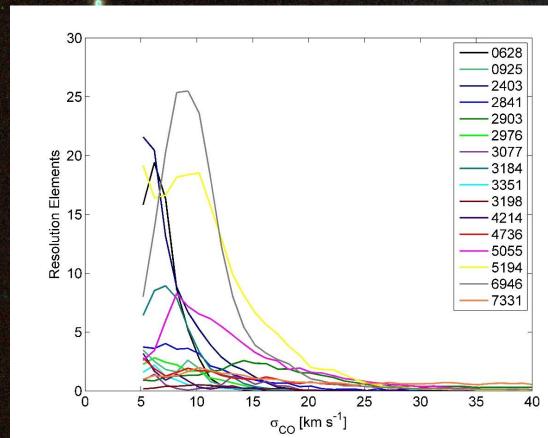
- Optical, NIR, UV, mm, Radio
- SF Dynamics:
  - Star formation processes, Disk & Cloud Stability, SF Law, Environmental Effects
    - THINGS+HERACLES+SINGS (Mogotsi+2016, Romeo & Mogotsi 2017, Caldu-Primo+), WIYN + SINGG + WISE (Mogotsi PhD thesis, Mogotsi+ in prep), ALMA
- SF Feedback:
  - Feedback: Winds & Outflows, Turbulence, Gas dynamics
    - SALT (see TALK), WiFeS & SINGG (see TALK), THINGS+HERACLES+SINGS (Mogotsi+2016, Mogotsi MSc thesis), MeerKAT (MHONGOOSE)
- SF Fuelling:
  - Gas accretion
    - MeerKAT (MHONGOOSE)

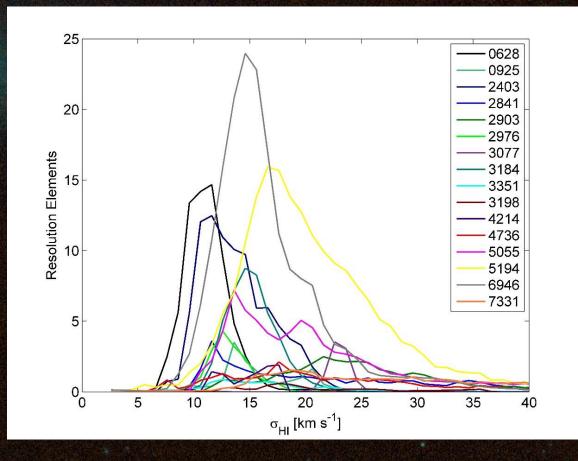
#### Star Formation



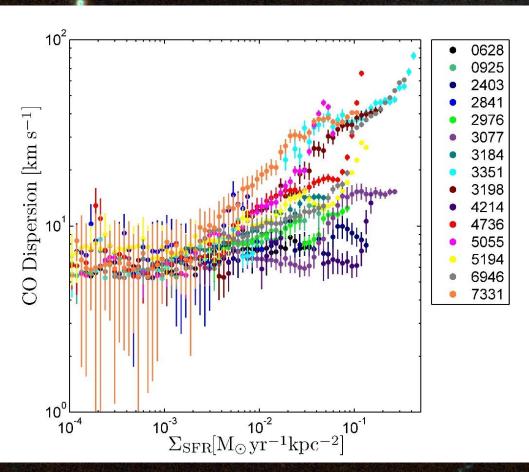
Spiral Arm Star Formation Sequence NASA / JPL-Caltech / M. Povich (Penn State Univ.) Spitzer Space Telescope • IRAC-MIPS sig10-009

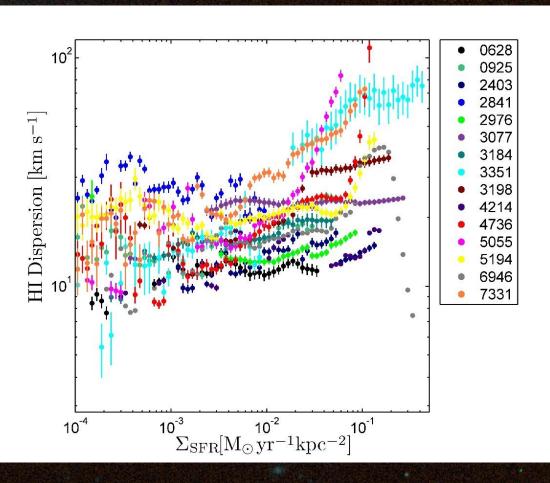
Mogotsi (MSc Thesis)





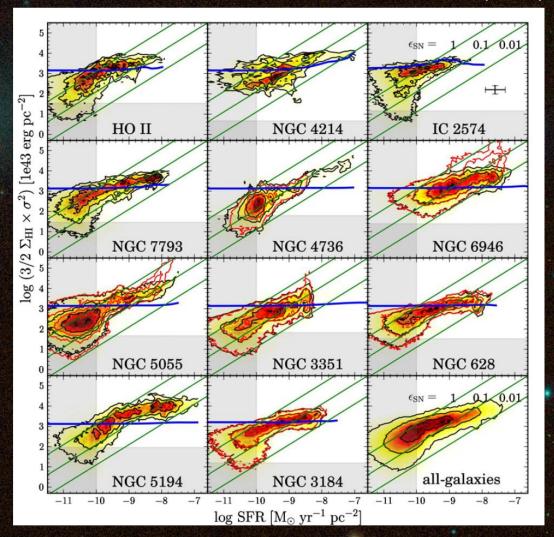
Mogotsi (MSc Thesis)





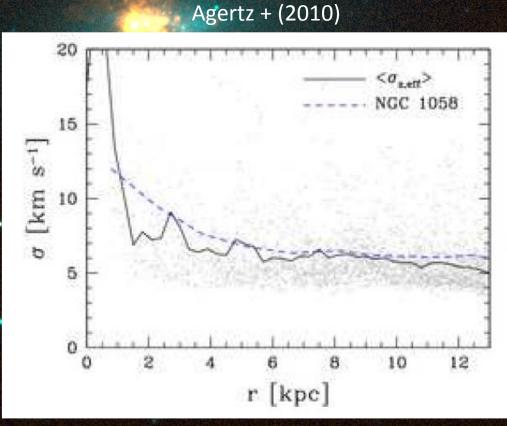
Tamburro + (2009)

- What drives the energetics of the ISM?
  - Does SF feedback explain all of it?
  - SN Feedback? Stellar Winds?
  - Magnetic Fields?

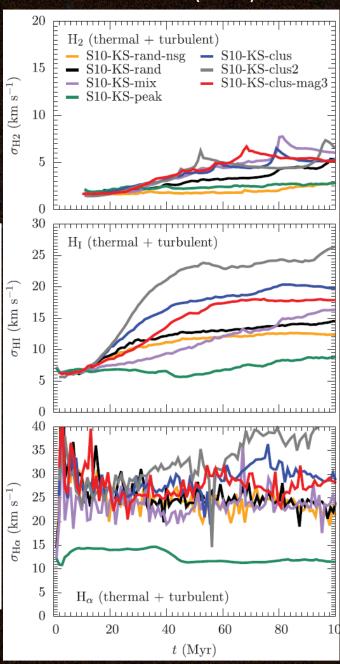


 What drives the energetics of the ISM?

- Does SF feedback...
- SN Feedback? Stellar Winds?
- Magnetic Fields?
- Test Simulations!



#### Girichidis + (2016)



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#### Feedback: Winds!

 Galactic, SN, stellar winds, outflows...

 Low redshift starbursts and SF galaxies can have much lower outflow velocities than LIRGS, high-z galaxies Schwartz & Martin 2004

Galaxy	v- $v$ <sub>sys</sub> (km s <sup>-1</sup> )
NGC 1569	-24
NGC 1614	-149
	+70
NGC 4214-2	-23
NGC 4449	-34
M82	-91
	-35
	+4
	+45
	+86
	,
NGC 2363	
NGC 4214-1	
NGC 5253	
I Zw 18	

Galaxy	$v$ - $v$ <sub>sys</sub> $(\text{km s}^{-1})$
NGC 1569	-24
NGC 1614	-149 +70

#### MHONGOOSE

- MeerKAT HI Observations of Nearby Galactic Objects: Observing Southern Emitters (de Blok+)
- How do galaxies get their gas?
- How is star formation regulated?
- How are outer disks & the cosmic web linked?
- Deep HI Observations
- 6000h 30 galaxies [originally]
- 1650h 30 galaxies [revised]
- MeerKAT
- SINGG-derived precursor sample (96 galaxies)
  - Sample (30) has been chosen
  - Ancillary Data, characterize final sample





### MHONGOOSE Update

- 6000h 30 galaxies [originally]
- 1650h 30 galaxies [revised]
- MeerKAT updated sensitivity
  - For HI to reach: 0.074mJy/beam [@5 km/s] needed 200hrs
  - Now to reach 0.074mJy/beam [@ 5 km/s] need 55hrs

#### SINGG

- Survey for Ionization in Neutral Gas Galaxies
  - Meurer et al. (2006),
  - Hanish et al. (2006),
  - UV subsample: Wong 2007
- H-alpha, R-band, UV photometry
- WISE NIR Data
  - (Jarrett priv. comm)

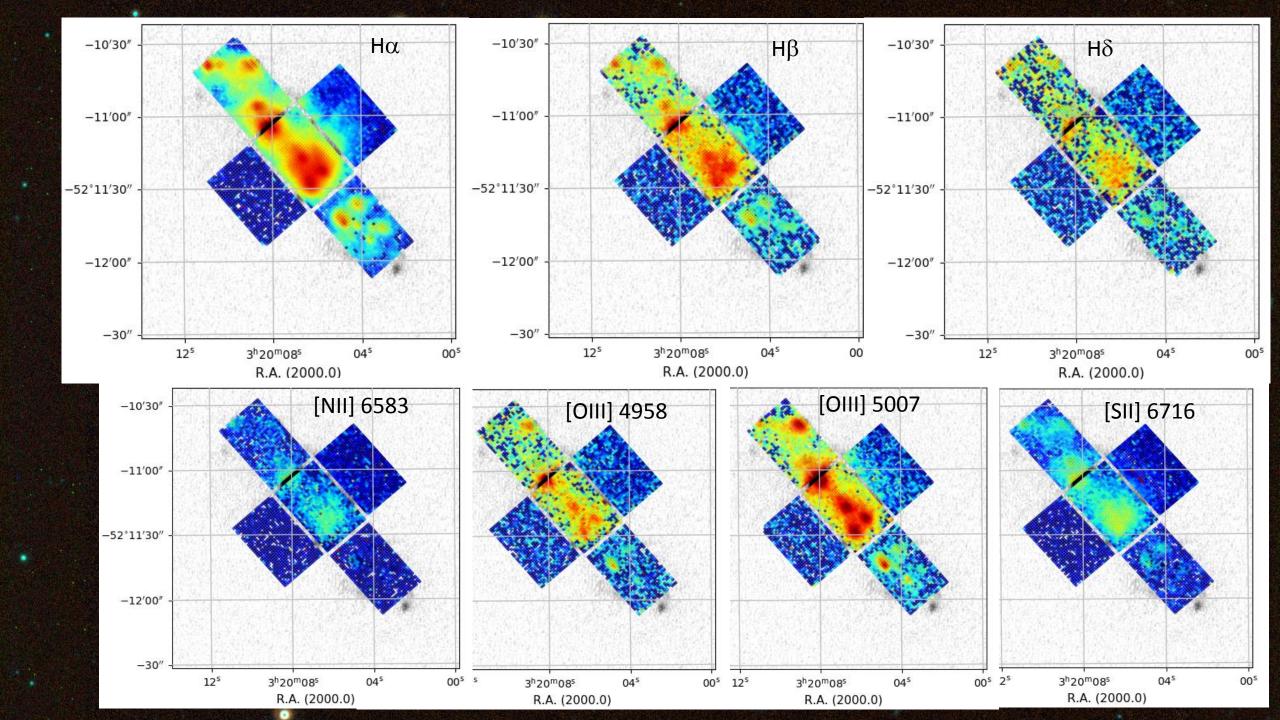
#### MHONGOOSE-WiFeS

with G. Meurer, + collaborators

- IFU Follow up of MHONGOOSE precursor sample
- Siding Spring: 2.3m ANU Telescope
- WiFeS: Dopita et al. (2007)
- IFU
- FOV: 25" x 38"
- 1" Res, Seeing-limited
- Blue: 320-590nm (0.077nm)
- Red: 530-706nm (0.044nm)

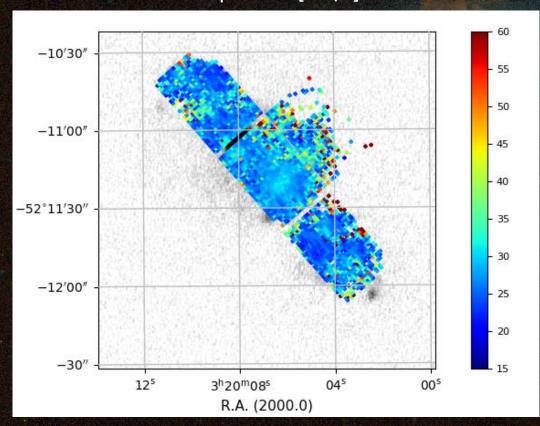
#### J0320-52

- NGC 1311
- D=6.99 Mpc
- Log MHI = 8.25
- Inc = 74 deg
- W50 = 30 km/s

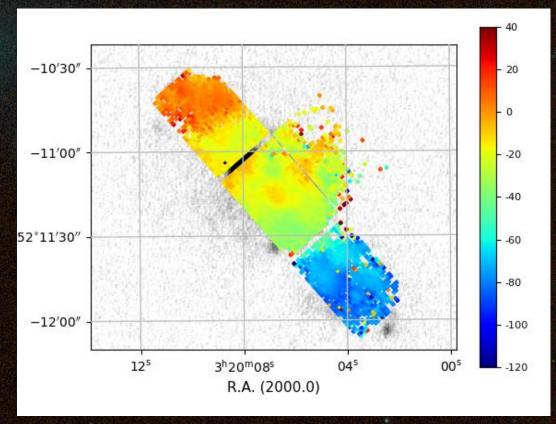


#### J0320-52 Kinematics

#### Dispersion [km/s]



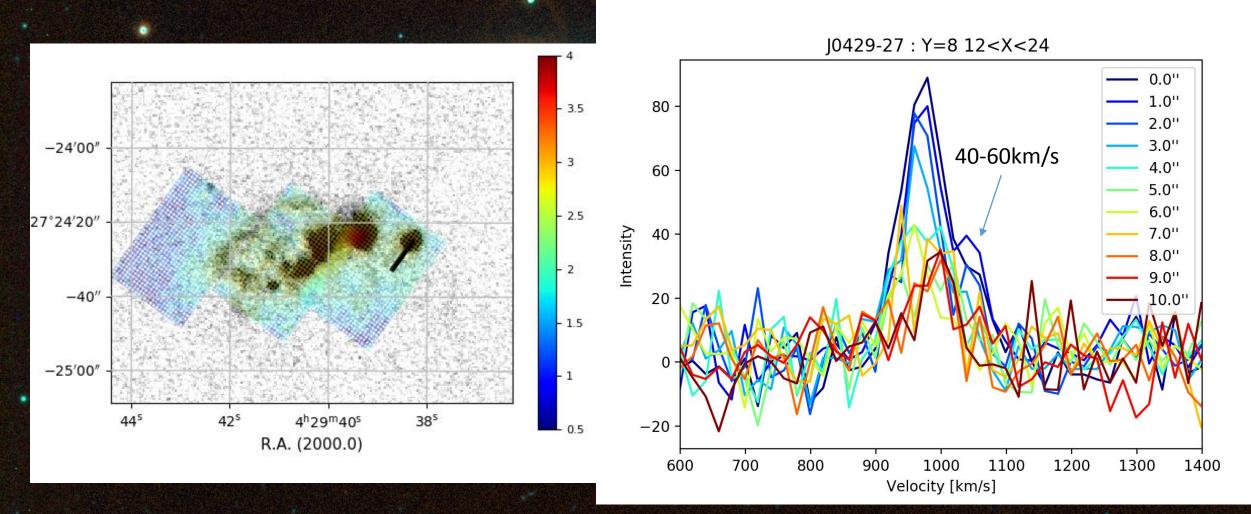
#### Velocity [km/s]



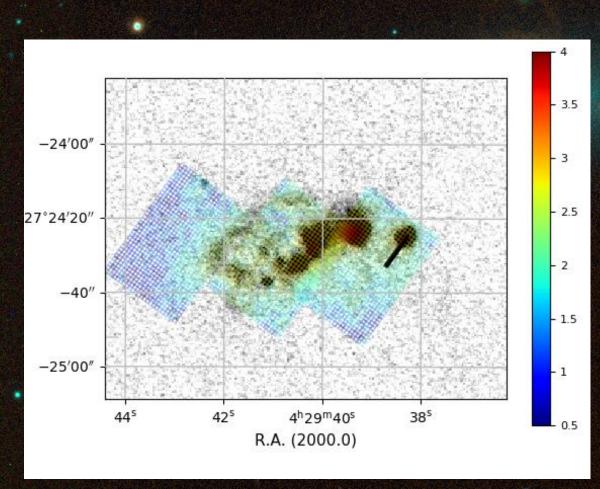
J0429-27

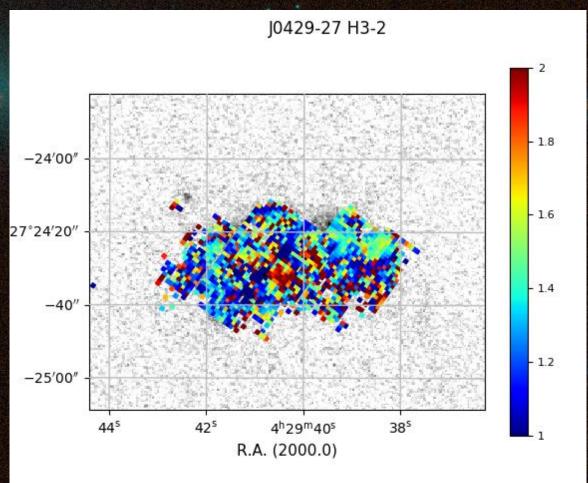
- NGC 1592
- D=13 Mpc
- Log MHI = 8.37
- Inc = 74 deg
- W50 = 73 km/s

### J0429-27 Winds? Profile Analysis

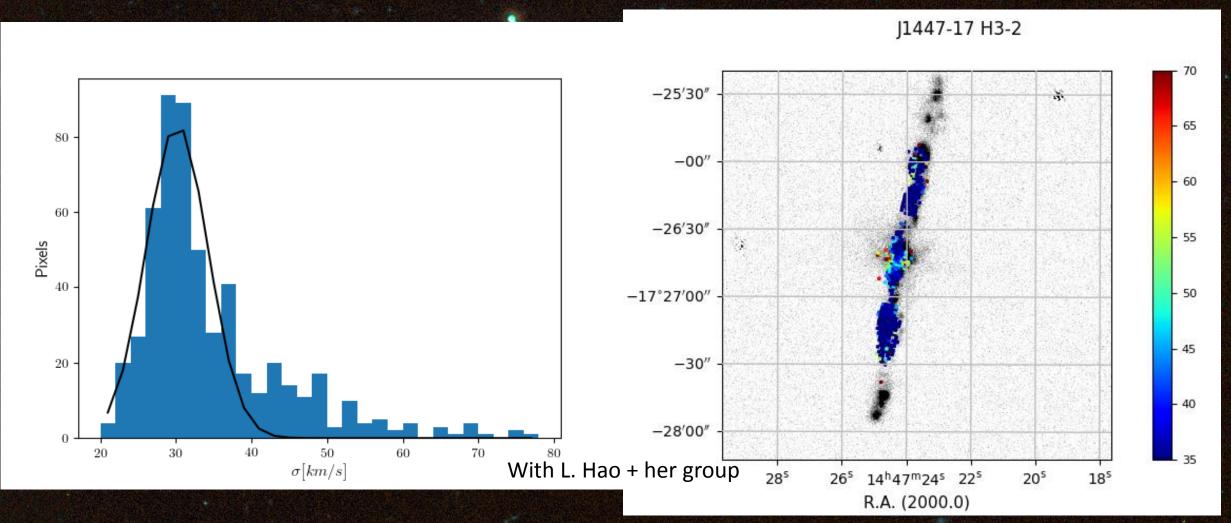


### J0429-27 Winds? Profile Asymmetry





### J1447-17 Winds? Dispersion Distributions



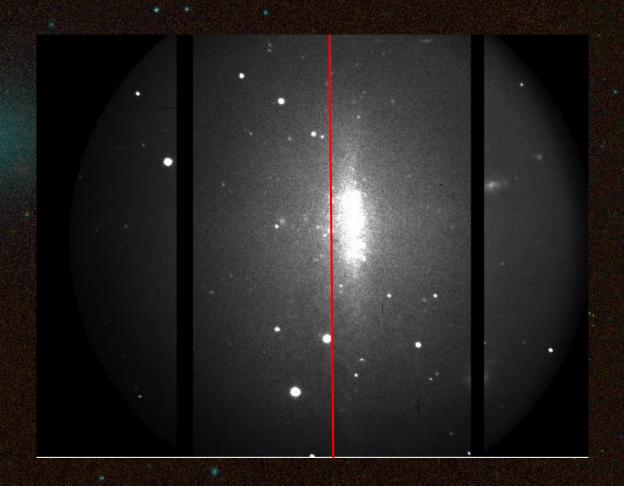
#### MHONGOOSE-WiFeS

- 4 Observing Runs
- >26 galaxies

- Stacking/smoothing and kinematic analysis
- ISM Characterization
- Stellar Population Analysis

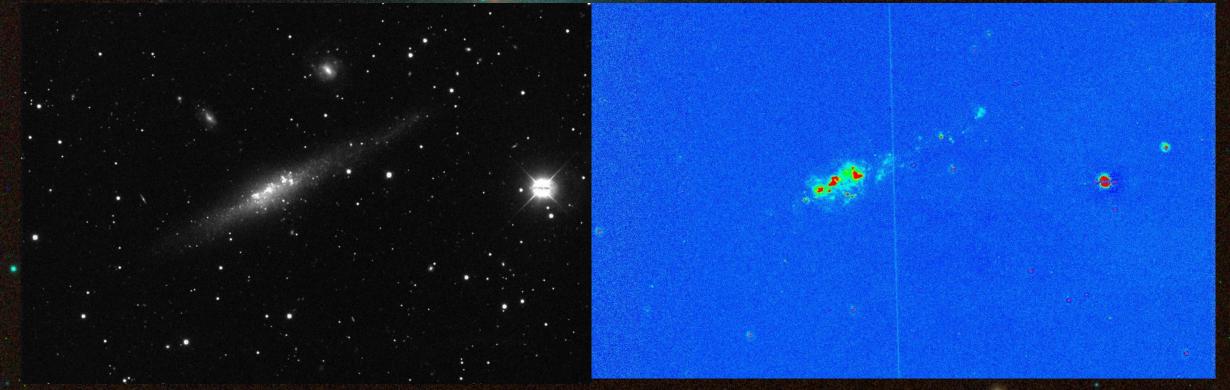
#### SALT - MHONGOOSE (with P. Väisänen)

- SALT RSS Long-slit Pilot Study
- Kinematics
- Outer Regions of Galaxies
- Diffuse Ionized Gas detection kinematics
- Winds/outflows



#### SALT WINDS & OUTFLOWS

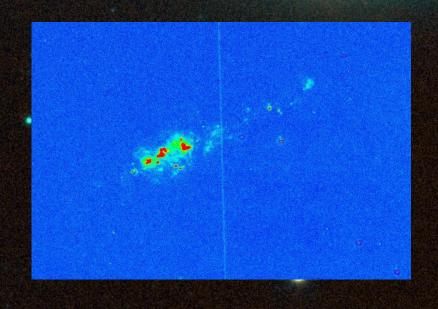
- J1303-17b (UGCA 320)
- D=7.74 Mpc
- Log (MHI) = 9.12 Msol

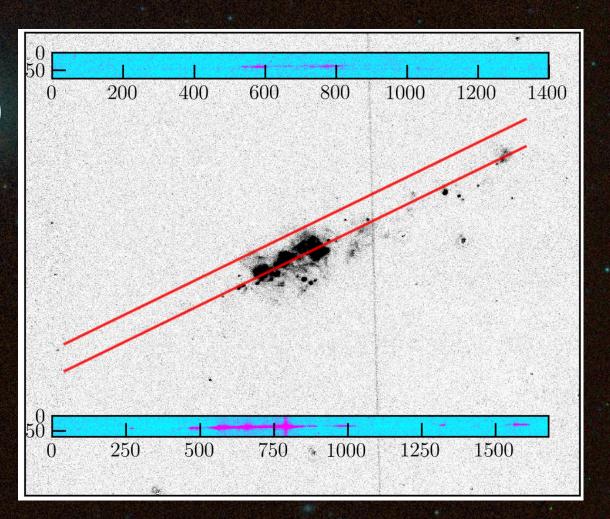


• 3 Slit Positions: Center + Off-plane

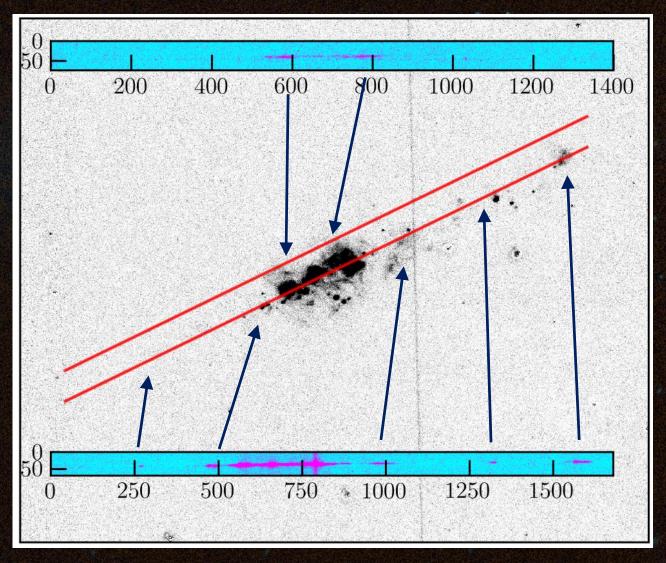
• PG2300 : Hα @ ~< R10,000

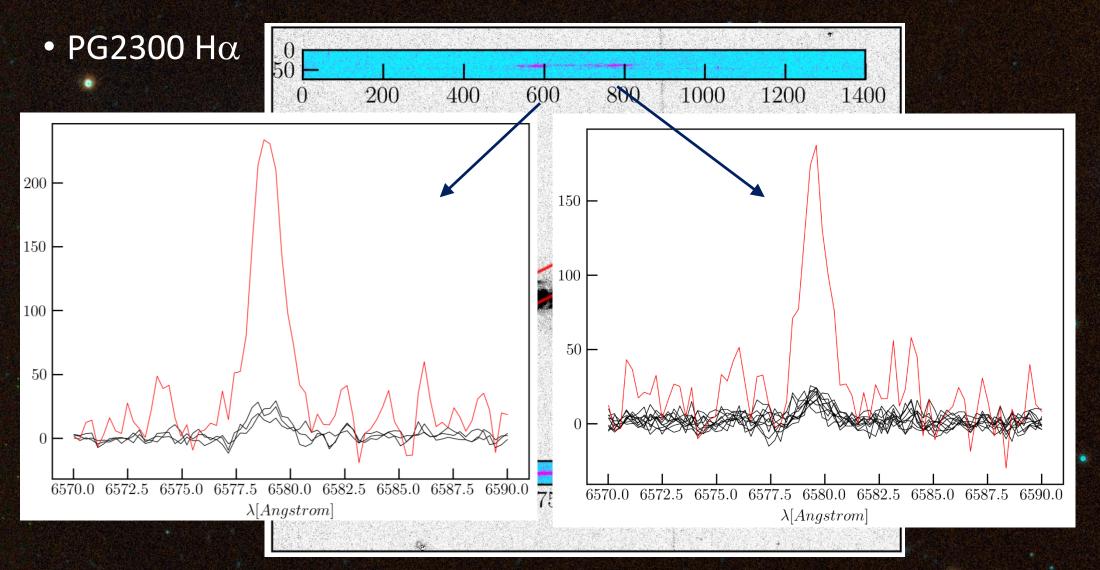
• PG1800 : Na D, SII,... @ R4,000

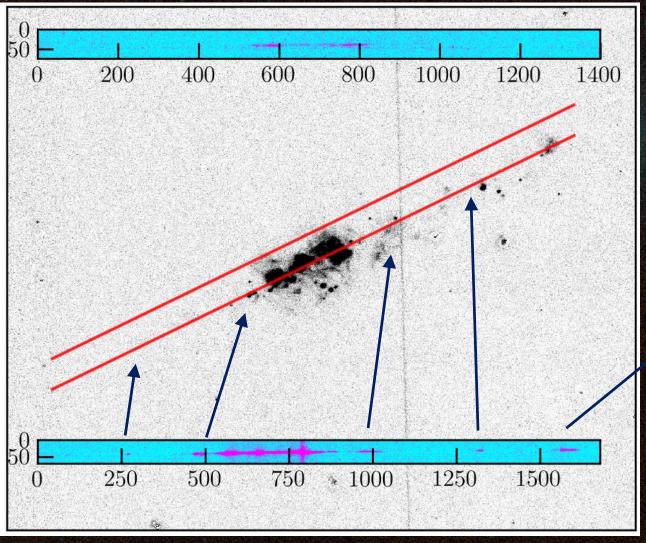


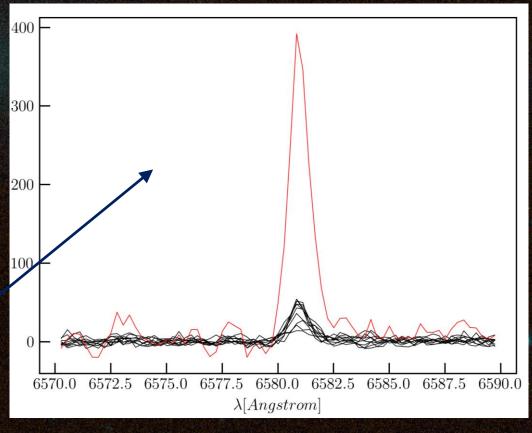


• PG2300 Hα









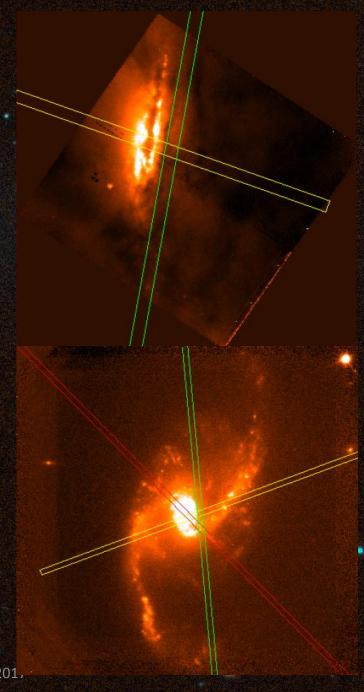
#### To LIRGS And Beyond!

(Väisänen, Rajin Ramphul, Melaku Sisay\* +)

SALT RSS Observations of LIRGS

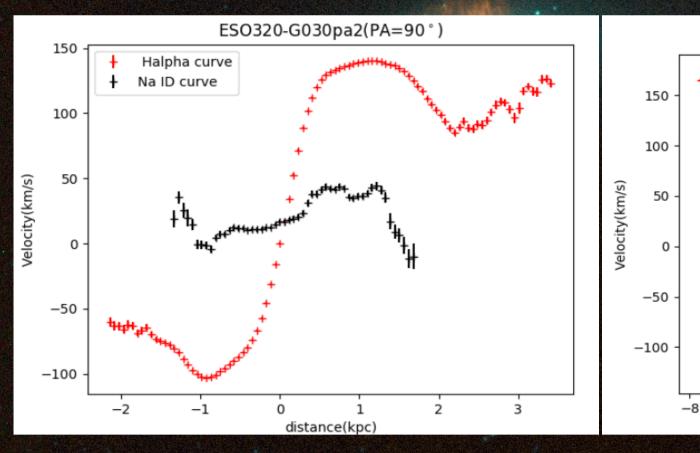
 Kinematics, Star Formation, Stellar Population Modelling & ISM Characterization

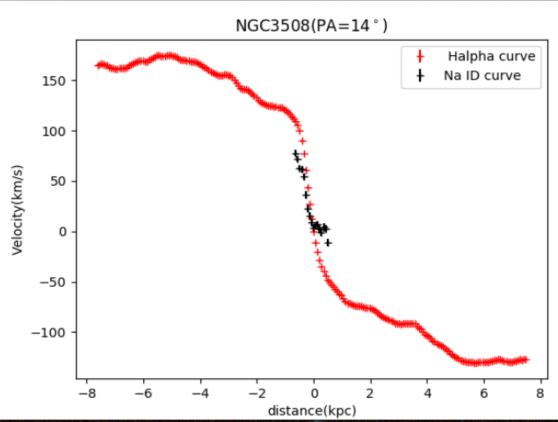
 Outflow Properties and Kinematics



#### To LIRGS And Beyond!

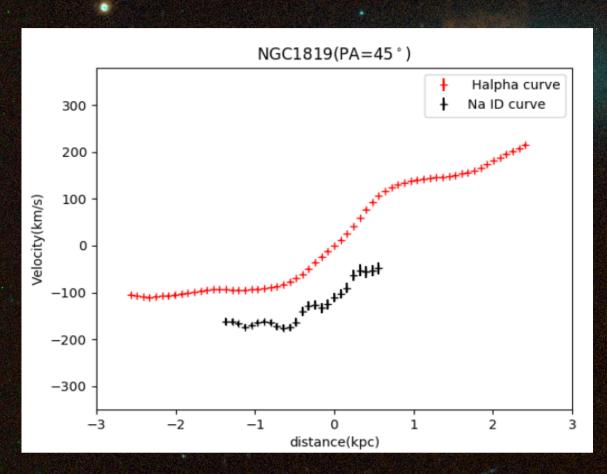
(Väisänen, Rajin Ramphul, Melaku Sisay\* +)

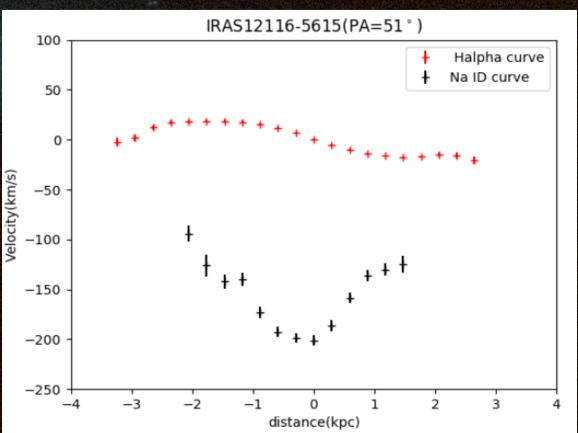




#### To LIRGS And Beyond!

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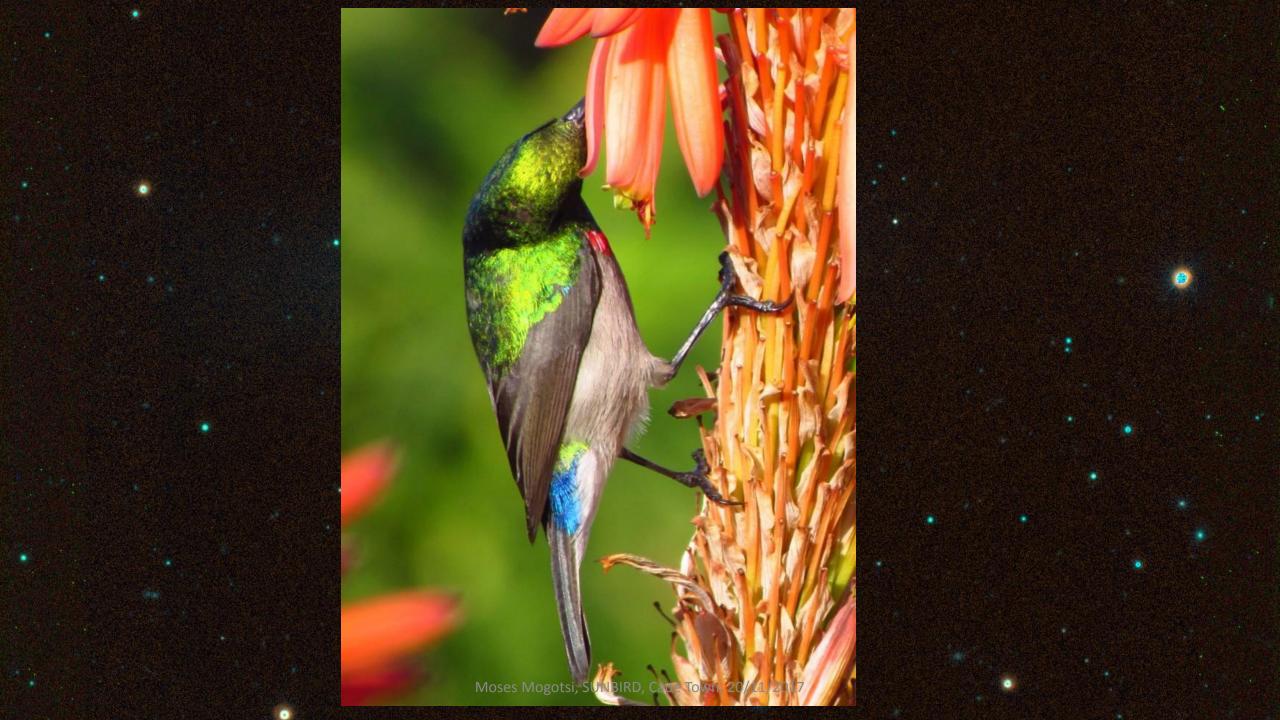




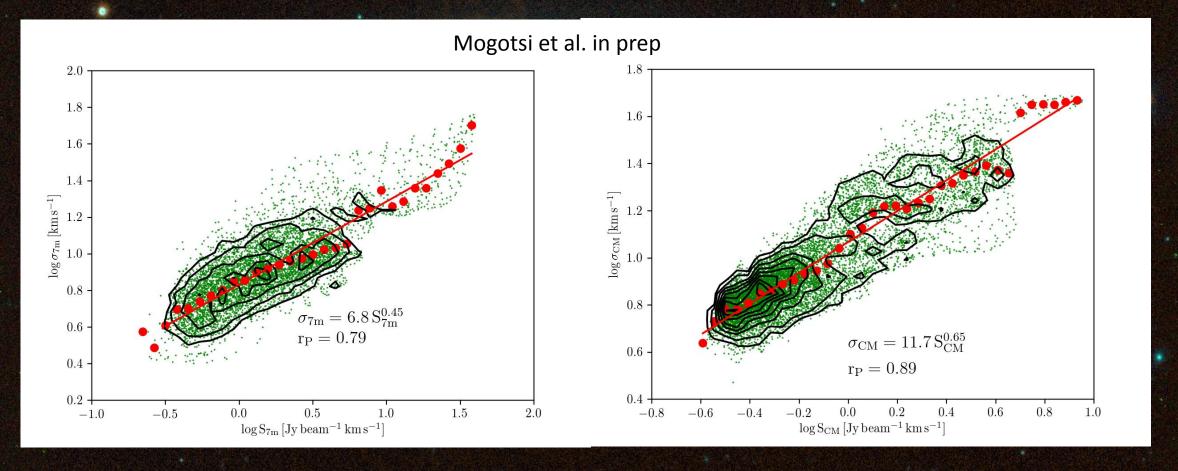
#### The Future

Further analysis of data

- SUNBIRD complement to MHONGOOSE?
  - Deep HI study of nearby starbursts and LIRGS
    - Environment diffuse gas in outer regions
    - Neutral gas kinematics, ouflow + ISM dynamics
    - Variability?



## Winds/Velocities via Gas Dispersions between GMCs at Large scales



#### The Bird

(Väisänen, Ramphul, Canizales-Romero and collaborators)

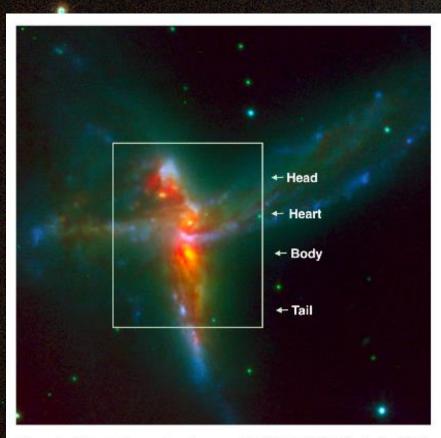
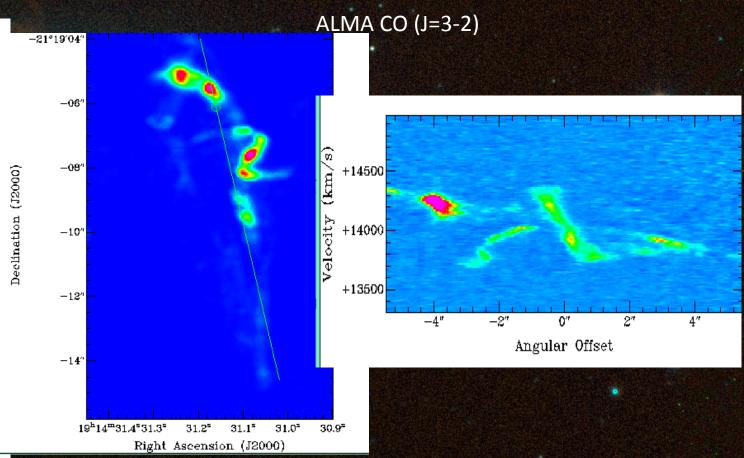


Figure 1. Left: A three-colour image of IRAS 19115-2124, aka the Bird, made by combining a K-band AO-image with HST I and B-bands (see V08). The box indicates the position of the SINFONI K250 datacube, it



Preliminary